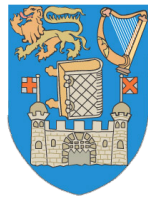


Introduction to Plasma Physics (PY5012)

Lecture 4: Single-Particle Motion



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Single-Particle Motion

- Motion in a uniform **B** field
 - **E** x **B** drift
 - Motion in non-uniform **B** fields
 - Gradient drift
 - Curvature drift
 - Other gradients of **B**
 - Time-varying E field
 - Time-varying B field
 - Adiabatic Invariants
- } Today's lecture

Drifts in uniform and non-uniform fields

- In last lecture, solved equations of motion for particles in uniform \mathbf{B} and \mathbf{E} fields.

- Showed that drift of guiding-centre is: $\mathbf{v}_E = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$

which can be extended to a form for a general force \mathbf{F} :

$$\mathbf{v}_f = \frac{1}{q} \frac{\mathbf{F} \times \mathbf{B}}{B^2}$$

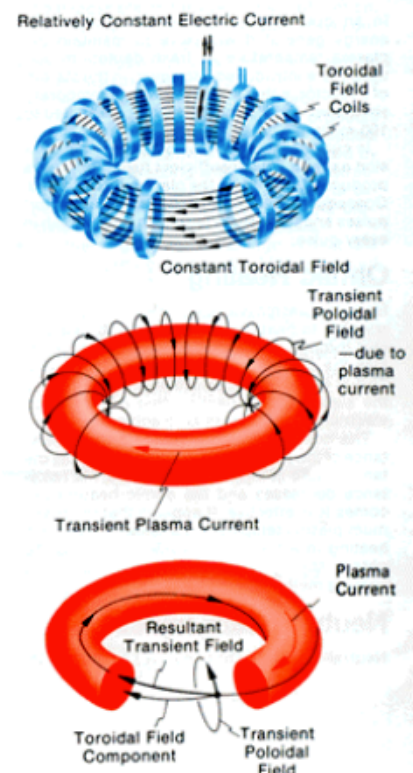
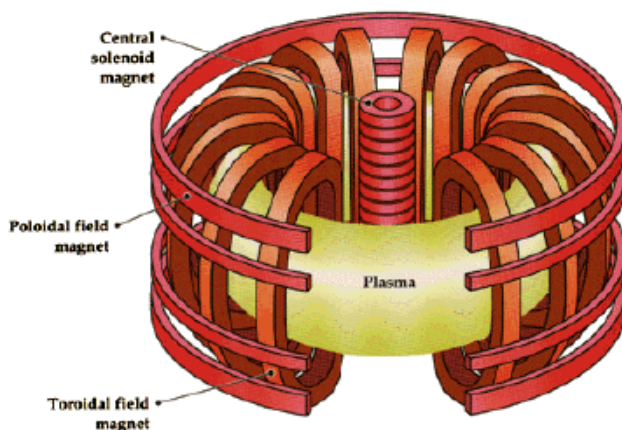
- In a gravitation field for example, $\mathbf{v}_g = \frac{m \mathbf{g} \times \mathbf{B}}{q B^2}$
- Similar to the drift \mathbf{v}_E , in that drift is perpendicular to both forces, but in this case particles of opposite charge drift in opposite direction.
- Uniform fields provide poor descriptions for many phenomena, such as planetary fields, coronal loops, tokamaks, which have spatially and temporally varying fields.

The configuration of a Tokamak.

TOroidalnaya KAmera ee MAgnitnaya Katushka,"

or "Toroidal Chamber in Magnetic Coil"

This is a device to magnetically confine a plasma, so that it can be heated to fusion temperatures.



Plasma confinement in Tokamak

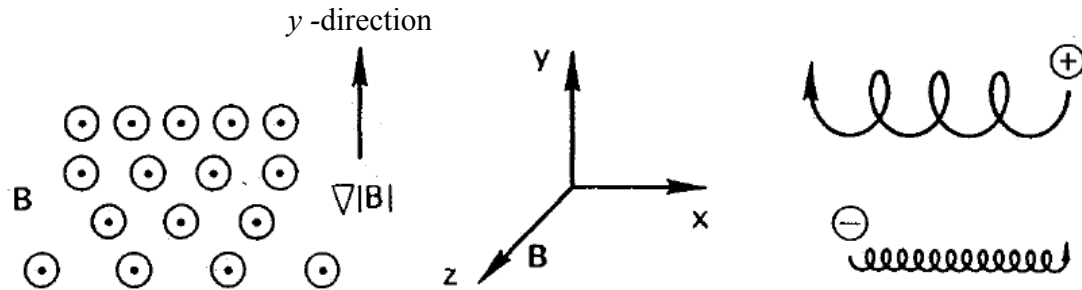


Drifts in non-uniform magnetic fields

- Particle drifts in inhomogeneous fields are classified in several ways.
- In this lecture, consider two drifts associated with spatially non-uniform \mathbf{B} : *gradient drift* and *curvature drift*. There are many others!
- As soon as introduce inhomogeneity, too complicated to obtain exact solutions for guiding centre drifts.
- Therefore use *orbit theory* approximation:
 - Within one Larmor orbit, \mathbf{B} is approximately uniform, i.e., the typical length-scale, L , over which \mathbf{B} varies is such that $L \gg r_L \Rightarrow$ gyro-orbit is nearly a circle.

Grad-B Drift

- Assume lines of forces are straight, but their density increases in y direction (see figure below from Chen, Page 27).



- Gradient in $|\mathbf{B}|$ causes the Larmor radius ($r_L = mv/qB$) to be larger at the bottom of the orbit than the top, which leads to a drift.
- Drift should be perpendicular to grad \mathbf{B} and \mathbf{B} and ions and electrons drift in opposite directions.

Grad-B Drift

- Consider spatially-varying magnetic field, $\mathbf{B} = (0, 0, B_z(y))$. i.e., \mathbf{B} only has z -component, the strength of which varies with y .

- Assume that $\mathbf{E} = 0$, so equation of motion is $\mathbf{F} = q(\mathbf{v} \times \mathbf{B})$

- Separating into components, $F_x = q(v_y B_z)$ (4.1a)

$$F_y = -q(v_x B_z) \quad (4.1b)$$

$$F_z = 0$$

- Now the gradient of B_z is $\frac{dB_z}{dy} \approx \frac{B_z}{L} \ll \frac{B_z}{r_L}$

$$\Rightarrow r_L \frac{dB_z}{dy} \ll B_z$$

- This means that the magnetic field strength can be expanded in a Taylor expansion for distances $y < r_L$:

$$B_z(y) = B_0 + y \frac{dB_z}{dy} + \dots$$

Grad-B Drift

- Expanding B_z to first order in Eqns. 4.1a and 4.1b:

$$F_x = qv_y \left(B_0 + y \frac{dB_z}{dy} \right) \quad (4.2)$$

$$F_y = -qv_x \left(B_0 + y \frac{dB_z}{dy} \right)$$

- Particles in \mathbf{B} -field travelling about a guide centre (0,0) have a helical trajectory:

$$x = r_L \sin(\omega_c t)$$

$$y = \pm r_L \cos(\omega_c t)$$

- The velocities can be written in a similar form: $v_x = v_\perp \cos(\omega_c t)$

$$v_y = \pm v_\perp \sin(\omega_c t)$$

- Substituting these into Eqn. 4.2 gives: $F_x = -qv_\perp \sin(\omega_c t) \left(B_0 \pm r_L \cos(\omega_c t) \frac{dB_z}{dy} \right)$

$$F_y = -qv_\perp \cos(\omega_c t) \left(B_0 \pm r_L \cos(\omega_c t) \frac{dB_z}{dy} \right)$$

Grad-B Drift

- Since we are only interested in the guiding centre motion, we average force over a gyroperiod. Therefore, in the x -direction:

$$\langle F_x \rangle = -qv_\perp \left[B_0 \langle \sin(\omega_c t) \rangle \pm r_L \langle \sin(\omega_c t) \cos(\omega_c t) \rangle \frac{dB_z}{dy} \right]$$

- But $\langle \sin(\omega_c t) \rangle = 0$ and $\langle \sin(\omega_c t) \cos(\omega_c t) \rangle = 0 \Rightarrow \langle F_x \rangle = 0$.

- In the y -direction: $\langle F_y \rangle = -qv_\perp \left[B_0 \langle \cos(\omega_c t) \rangle \pm r_L \langle \cos^2(\omega_c t) \rangle \frac{dB_z}{dy} \right]$

$$= \mp \frac{qv_\perp r_L}{2} \frac{dB_z}{dy} \quad (4.3)$$

where $\langle \cos(\omega_c t) \rangle = 0$ but $\langle \cos^2(\omega_c t) \rangle = 1/2$.

Grad-B Drift

- In general, drift of guiding-centre is $\mathbf{v}_f = \frac{1}{q} \frac{\mathbf{F} \times \mathbf{B}}{B^2}$

- Subing in using Eqn 4.3:

$$\begin{aligned} \mathbf{v}_{\nabla B} &= \frac{1}{q} \frac{\langle F_y \rangle \hat{\mathbf{y}} \times \mathbf{B}_z \hat{\mathbf{z}}}{B_z^2} \\ &= \mp \frac{v_{\perp} r_L}{2B_z} \frac{dB_z}{dy} \hat{\mathbf{x}} \end{aligned}$$

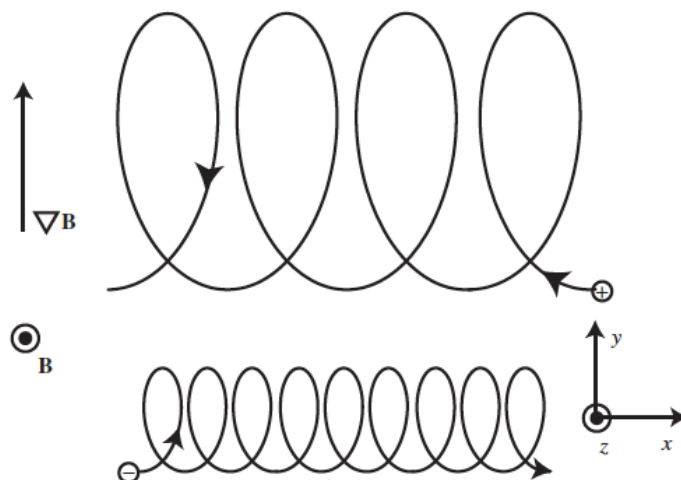
- So +ve particles drift in $-x$ direction and $-ve$ particles drift in $+x$.
- In 3D, the result can be generalised to:

$$\mathbf{v}_{\nabla B} = \pm \frac{1}{2} v_{\perp} r_L \frac{\mathbf{B} \times \nabla B}{B^2} \quad \text{Grad-B Drift}$$

- The \pm stands for sign of charge. The grad-B drift is in opposite directions for electrons and ions and causes a current transverse to \mathbf{B} .

Grad-B Drift

- Below are shown particle drifts due to a magnetic field gradient, where $\mathbf{B}(y) = z B_z(y)$. Page 31 of Inan & Golkowski.



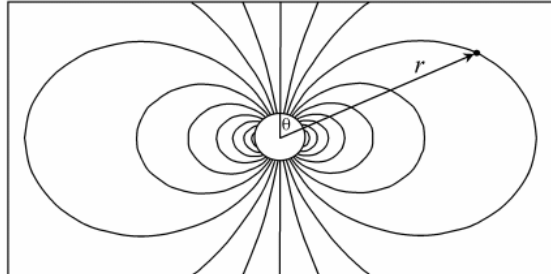
- Consider $r_L = mv/qB$. Local gyroradius is large where \mathbf{B} is small and visa versa which gives rise to a drift. Note also that charge determines direction of drift.

Grad-B Drift: Planetary Ring Current

- Gradient drift is responsible for current in inner parts of planetary magnetospheres.
- Approximate field is a dipole:

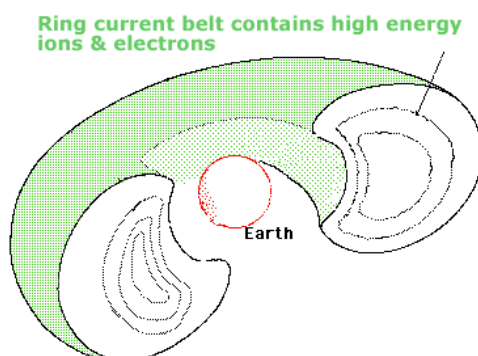
$$B_r = \frac{\mu_0}{4\pi} \frac{2m}{r^3} \cos(\theta)$$

$$B_\theta = \frac{\mu_0}{4\pi} \frac{2m}{r^3} \sin(\theta)$$

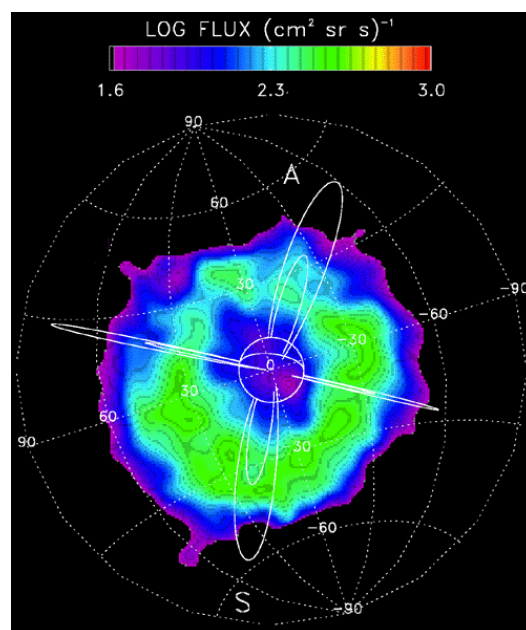


- In equatorial plane, $B_r = 0$ and $B_\theta = B_\theta / r^3$. There is therefore a positive gradient in B_θ directed radially outwards.
- There is therefore a grad-B drift perpendicular to \mathbf{B} and grad-B, which produces a ring current circulating about a planet.

Grad-B Drift: Planetary Ring Current



- Right: Ring current viewed from north pole with NASA's Image satellite.



Curvature Drift

- When charged particles move along curved magnetic field lines, experience centrifugal force perpendicular to magnetic field (Fig Inan & Golkowski, Page 34).

- Assume radius of curvature (R_c) is $\gg r_L$

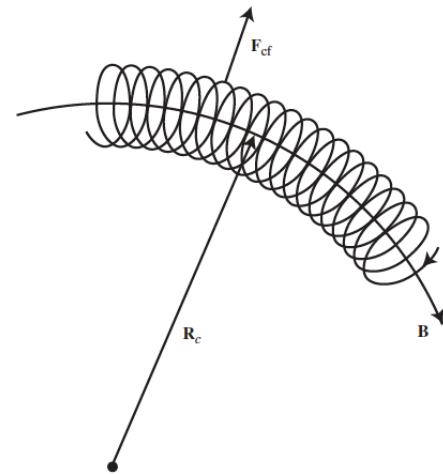
- The outward centrifugal force is

$$\mathbf{F}_{cf} = \frac{mv_{\parallel}^2}{R_c} \hat{\mathbf{r}}$$

- This can be directly inserted into general form for guiding-centre drift

$$\mathbf{v}_f = \frac{1}{q} \frac{\mathbf{F} \times \mathbf{B}}{B^2}$$

$$\Rightarrow \mathbf{v}_R = \frac{mv_{\parallel}^2}{qR_c^2} \frac{\mathbf{R}_c \times \mathbf{B}}{B^2}$$



Curvature Drift

- Drift is therefore into or out of page depending on sign of q .

Curvature and grad-B drift

- Type of field configuration studied above – having curved but parallel field lines – will never occur in reality, since $\nabla \cdot \mathbf{B} \neq 0$ for this field.

- In practice, curved field lines will always be converging/diverging. Thus a curvature drift will always be accompanied by a grad-B drift.

- Where ∇B is in the opposite direction to R_c , the grad-B and curvature drifts act in the same direction.

- Right is a cross-section of a tokamak magnetic field.

- Dashed lines are field lines. In the highlighted section, the field is lower at edge than in the centre. So curvature and gradient drift act in same direction - outwards for ions – so ions hit the vessel walls and are lost.

