

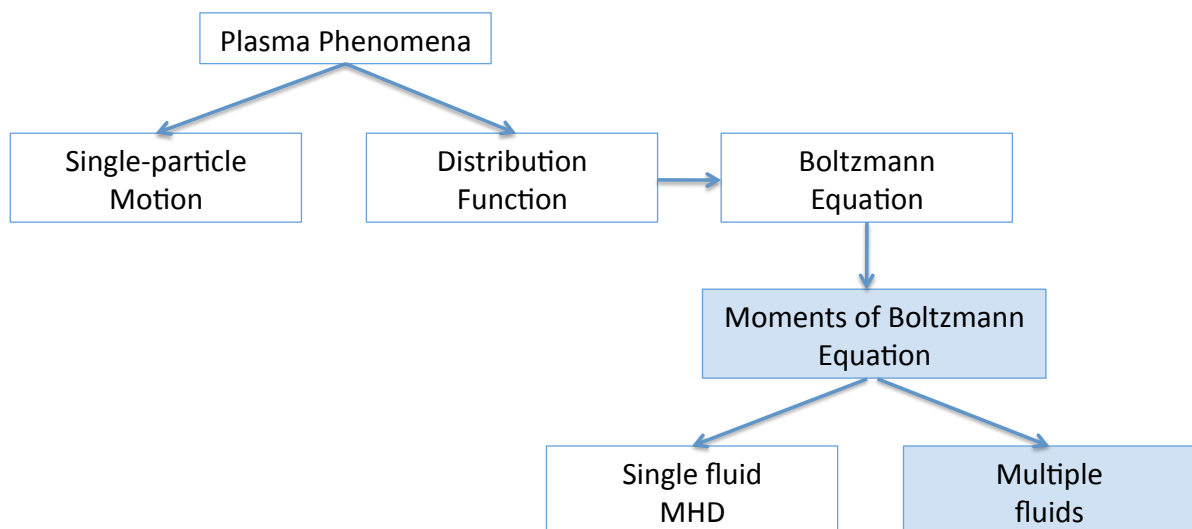
Introduction to Plasma Physics (PY5012)

Lecture 8: Multiple-fluid theory of plasmas



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Hierarchy of plasma phenomena



Fluid Approach to Plasmas

- Fluid approach describes bulk properties of plasma. We do not attempt to solve unique trajectories of all particles in a plasma. This simplification works very well for majority of plasmas, despite gross simplifications made.
- Fluid theory follows directly from moments of the Boltzmann equation (Lecture 7).
- Each of the moments of the Boltzmann equation is a transport equation describing the dynamics of a quantity associated with a given power of \mathbf{v} .

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{u}) = 0 \quad \text{Continuity of mass or charge transport}$$

$$mn \left[\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} \right] = qn(\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla \cdot \mathbf{P} + \mathbf{P}_{ij} \quad \text{Momentum Transport}$$

$$\frac{\partial}{\partial t} \left[n \frac{1}{2} m u^2 \right] + \nabla \cdot \left[n \frac{1}{2} m \langle u^2 \mathbf{u} \rangle \right] - nq \langle \mathbf{E} \cdot \mathbf{u} \rangle = \frac{m}{2} \int u^2 \left(\frac{\partial f}{\partial t} \right)_{coll} d\mathbf{u} \quad \text{Energy Transport}$$

Cold-Plasma Model

- Simplest set of macroscopic equations can be obtained by simplifying the momentum transfer equation and neglect thermal motions of particles.
- Here, set kinetic pressure tensor to zero, i.e., $\mathbf{P} = mn \langle \mathbf{w}\mathbf{w} \rangle = 0$ as $\mathbf{w} = 0$.
- Remaining macroscopic variables are then n and \mathbf{u} , described by

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{u}) = 0$$

$$mn \left[\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} \right] = qn(\mathbf{E} + \mathbf{u} \times \mathbf{B}) + \mathbf{P}_{ij}$$

- Collision term \mathbf{P}_{ij} can be approximated by by an “effective” collision frequency.
- Assumed that collisions cause a rate of decrease in momentum:

$$\mathbf{P}_{ij} = -mn\nu_{eff} \mathbf{u}$$

Warm-Plasma Model

- Alternative set of macroscopic equation is obtained by truncating energy conservation equation.

- Consider pressure tensor: $\mathbf{P} = mn\langle\mathbf{w}\mathbf{w}\rangle = \begin{pmatrix} P_{xx} & P_{xy} & P_{xz} \\ P_{yx} & P_{yy} & P_{yz} \\ P_{zx} & P_{zy} & P_{zz} \end{pmatrix}$

- Components represent transport of momentum. Diagonal elements represent pressure, while off-diagonal represent shearing stresses.

- In warm-plasma model, only consider diagonal pressure elements, so

$$\nabla \cdot \mathbf{P} = \nabla p$$

- That is, viscous forces are neglected. We then have

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{u}) = 0$$

$$mn \left[\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u} \right] = qn(\mathbf{E} + \mathbf{u} \times \mathbf{B}) - \nabla p + \mathbf{P}_{ij}$$

Warm-Plasma Model

- The previous system of equations does not form a closed set, since scalar pressure is now a third variable. Usually determined by energy equation

- If plasma is *isothermal*, assume *equation of state* of form:

$$p = nk_B T \quad \text{and} \quad \nabla p = k_B T \nabla n$$

- Holds for slow time variations, allowing temperatures to reach equilibrium.

- If plasma does not exchange energy with its surrounds, assume it is *adiabatic*:

$$p n^\gamma = \text{constant} \quad \text{and} \quad \frac{\nabla p}{p} = \gamma \frac{\nabla n}{n}$$

where γ is the ratio of the specific heats at constant pressure.

Simplified Energy Equation

- Note, the energy equation can be written

$$\frac{\partial[1/2nm \langle w^2 \rangle]}{\partial t} + \nabla \cdot (1/2nm \langle w^2 \rangle \mathbf{u}) + (\mathbf{P} \cdot \nabla) \mathbf{u} + \nabla \cdot \mathbf{q} = P_{ij}$$

where \mathbf{q} is the heat flow vector. For electrons, commonly used approximation for \mathbf{q} is

$$\mathbf{q} = K \nabla T$$

where K is the thermal *Spitzer conductivity*.

- As average energy of plasma is $1/2m \langle \mathbf{w}\mathbf{w} \rangle = 3/2 k_B T$ and using $p = n k_B T$
 $\Rightarrow 3/2 p = 1/2nm \langle \mathbf{w}\mathbf{w} \rangle$. Energy equation can then be written

$$\frac{\partial(3/2p)}{\partial t} + \nabla \cdot (3/2p\mathbf{u}) - p \nabla \cdot \mathbf{u} + \nabla \cdot \mathbf{q} = P_{ij}$$

- The quantity $3/2p\mathbf{u}$ represents the flow of energy density at the fluid velocity.

Complete set of two-fluid equations

- Consider plasma of two species; ions and electrons, in which fluid is fully ionised, isotropic and collisionless. The charge and current densities are

$$\sigma = n_i q_i + n_e q_e$$

$$\mathbf{j} = n_i q_i \mathbf{v}_i + n_e q_e \mathbf{v}_e$$

- Using $\mathbf{v} = \mathbf{u}$, complete set of electrodynamics equations are then ($j = i$ or e)

$$m_j n_j \left[\frac{\partial \mathbf{v}_j}{\partial t} + (\mathbf{v}_j \cdot \nabla) \mathbf{v}_j \right] = -\nabla p_j - q_j n_j (\mathbf{E} + \mathbf{v}_j \times \mathbf{B})$$

$$\frac{\partial n_j}{\partial t} + \nabla \cdot (n_j \mathbf{v}_j) = 0$$

$$\epsilon_0 \nabla \cdot \mathbf{E} = \sigma$$

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t}$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\frac{1}{\mu_0} \nabla \times \mathbf{B} = \mathbf{j} + \epsilon_0 \frac{\partial \mathbf{E}}{\partial t}$$

$$p_j = C_j n_j^{\gamma_j}$$

Fluids drifts perpendicular to \mathbf{B}

- Since a fluid element is composed of many individual particles, expect drifts perpendicular to \mathbf{B} . But, the $\text{grad}(p)$ term results in a fluid drift called *diamagnetic drift*.

- Consider momentum equation for each species:

$$mn \left[\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \right] = -\nabla p - qn(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

$$(1) \quad (2) \quad (3)$$

- Consider ratio of terms (1) to (3):

$$\frac{(1)}{(3)} \approx \left| \frac{mni\omega_{\perp}}{qnv_{\perp}B} \right| \approx \frac{\omega}{\omega_c}$$

- Here we have used $\partial/\partial t = i\omega$. If only consider slow drifts compared to time-scale of the gyrofrequency, can set (1) to zero.

Fluids drifts perpendicular to \mathbf{B}

- Therefore can write $0 \approx -qn(\mathbf{E} + \mathbf{v}_{\perp} \times \mathbf{B}) - \nabla p$

$$\text{where } \mathbf{v} \times \mathbf{B} = (\mathbf{v}_{\perp} + \mathbf{v}_{\parallel}) \times \mathbf{B} = \mathbf{v}_{\perp} \times \mathbf{B}$$

- Taking the cross-product $0 = qn[\mathbf{E} \times \mathbf{B} + (\mathbf{v}_{\perp} \times \mathbf{B}) \times \mathbf{B}] - \nabla p \times \mathbf{B}$

- Using the identity $(\mathbf{A} \times \mathbf{B}) \times \mathbf{C} = (\mathbf{C} \cdot \mathbf{A})\mathbf{B} - (\mathbf{C} \cdot \mathbf{B})\mathbf{A}$ we can write

$$0 = qn[\mathbf{E} \times \mathbf{B} + (\mathbf{v}_{\perp} \cdot \mathbf{B})\mathbf{B} - (\mathbf{B} \cdot \mathbf{B})\mathbf{v}_{\perp}] - \nabla p \times \mathbf{B}$$

- As \mathbf{v}_{\perp} is perpendicular to \mathbf{B} , $\mathbf{v}_{\perp} \cdot \mathbf{B} = 0$. Therefore

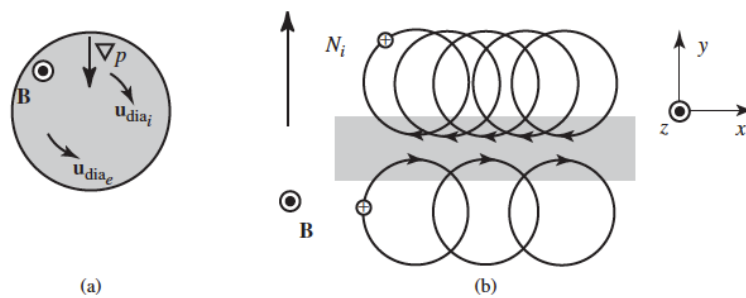
$$\begin{aligned} \mathbf{v}_{\perp} &= \frac{\mathbf{E} \times \mathbf{B}}{B^2} - \frac{\nabla p \times \mathbf{B}}{qnB^2} \\ &= \mathbf{v}_E + \mathbf{v}_D \end{aligned}$$

Fluids drifts perpendicular to B

- In previous equation is $\mathbf{v}_E = \frac{\mathbf{E} \times \mathbf{B}}{B^2}$ $\mathbf{E} \times \mathbf{B}$ drift

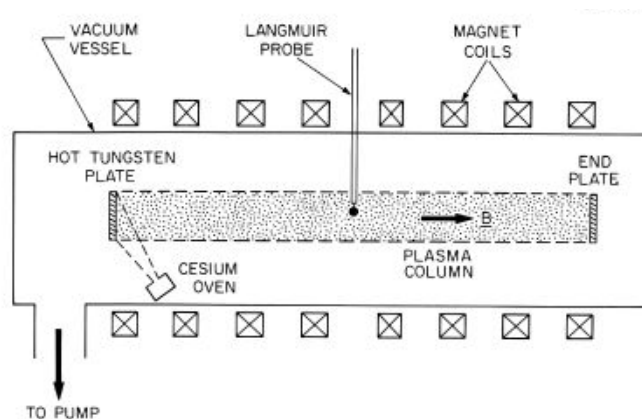
and $\mathbf{v}_D = -\frac{\nabla p \times \mathbf{B}}{qnB^2}$ is *diamagnetic drift*.

- The \mathbf{v}_E drift is same as for guiding centres, but there is now a new drift, called the diamagnetic drift. Is in opposite directions for ions and electrons.
- Gives currents in plasma that reduce magnetic field in plasma. More ions moving to left in shaded area that to the right (Inan & Golkowski, Page 111).



Diamagnetic Drift in Q-machines

- Diamagnetic drift first measured in *Q*-machines



- See <http://www.physics.uiowa.edu/xplasma/Qmachine.html>

Fluid drifts parallel to \mathbf{B}

- Consider z component of fluid equation of motion:

$$mn \left[\frac{\partial v_z}{\partial t} + (\mathbf{v} \cdot \nabla) v_z \right] = -qnE_z - \frac{\partial p}{\partial z}$$

- The convective term can be neglected as it is much smaller than $\partial v_z / \partial t$
- Using $p = n k_B T$ or $\frac{\partial p}{\partial z} = k_B T \frac{\partial n}{\partial z}$ we can write

$$\frac{\partial v_z}{\partial t} = \frac{q}{m} E_z - \frac{\gamma k_B T}{mn} \frac{\partial n}{\partial z}$$

- This shows that the fluid is accelerated along \mathbf{B} under the combined electrostatic and pressure gradient forces.

Fluid drifts parallel to \mathbf{B}

- Taking the limit as $m \rightarrow 0$ and $q = -e$ and $E_z = -\partial\phi/dz$ we have

$$eE_z = e \frac{\partial\phi}{dz} = \frac{\gamma k_B T}{n} \frac{\partial n}{dz}$$

- Electrons are so mobile that their heat conductivity is almost infinite.
- Assuming isothermal electrons and taking $\gamma = 1$, we can integrate to get

$$e\phi = k_B T \ln(n) + C$$

- We can therefore write

$$n = n_0 \exp(e\phi/k_B T)$$

- This is called the *Boltzmann relation* or *Boltzmann factor* for electrons.
- Implies that electrons have a tendency to move rapidly in response to an external force (i.e., electrostatic potential gradient).